

Hiragino Kaku Gothic Std W8 31

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we have again identified three groups of
candidates (Figure \[fig:mips\]) which are
all relatively well-matched with our
model, and which we here term the first
three tiers of candidates. ![Magnitude
versus the predicted \$Y\$ band absolute
magnitude for white dwarf candidates
from the MPFC15 White Dwarf Catalog
with no reliable constraint on parallax
(but observed to be within \$2\farcs4\$ of a
star to form a close pair). []{data-label="f
ig:mips"}](f8_color.pdf){width="50.0000
0%"} Conclusion ===== We have
presented the first up-to-date (April

2018) white dwarf catalog from the MPC.

We also presented the first model of the absolute magnitudes of white dwarfs as a function of temperature, and explain why the model is so close to the empirical

Johnson-Kron-Cousins system. Our model allows us to calculate the distance that each white dwarf will have based on its absolute I magnitude, and we put it online at [We have also explained how to

use this model to quickly and easily identify the right WDs for spectroscopy at high proper motion. Bailer-Jones, C. A.

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